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STEP AUTHORS:

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TITLE:

Increase of radiation intensity before a shock wave reaches the surface of a homogeneous medium

PERIODICAL: Ukrayins'kyy fizychnyy zhurnal, v. 7, no. 10, 1962,

1083-1088

TEXT: It is assumed that the shock wave moves with constant velocity and the energy F radiated from a surface unit of the front is independent of time. The amount of energy J(u) radiated from  $u = -\infty$  to u, when the wave reaches an optical depth  $(u_0 - u)v$ , is computed by integrating the probability function given by V. V. Sobolev;  $u = t/t_1$ ,  $t_1$  being the average duration of a quantum in absorbed state; v is the dimensionless velocity of the wave:

Card 1/3

$$J(u) = \frac{F}{v^{2}(1 - a^{2})} \left[ Q_{1}e^{-(u_{0}-u)v\sqrt{\alpha_{1}}} + Q_{2}e^{-(u_{0}-u)v\sqrt{\alpha_{2}}} + Q_{3}e^{-(u_{0}-u)v\sqrt{\alpha_{3}}} \right]$$
(9)

Here  $a^2 = 1 - \lambda$ ,  $\lambda$  is the relative amount of scattered energy,  $\alpha_1$ ,  $\alpha_2$ ,  $\alpha_3$  are the roots of

$$(a^{2} + x^{2})^{2} + v^{2}x^{2}(1 + x^{2})^{2} = (x^{2} + \alpha_{1})(x^{2} + \alpha_{2})(x^{2} + \alpha_{3})$$
 (7)

 $Q_1$ ,  $Q_2$ ,  $Q_3$  depend on these roots, and on v and a. Simplified expressions are obtained for several special cases. If the velocity Card 2/3

is 100 km/sec and the density  $n=10^{12}~\rm cm^{-3}$  brightness increases e times in  $10^{-5}$  sec near spectral line frequency. The above expression does not apply to rarefied media if the frequency is near that of a spectral line. The authors thank Professor S. A. Kaplan for the formulation of the subject of the paper.

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